

37.4 Events and Measurements

To question some of our most basic assumptions about space and time requires extreme care. We need to be certain that no assumptions slip into our analysis unnoticed. Our goal is to describe the motion of a particle in a clear and precise way, making the barest minimum of assumptions.

Events

The fundamental entity of relativity is called an **event**. An event is a physical activity that takes place at a definite point in space and at a definite instant of time. An exploding firecracker is an event. A collision between two particles is an event. A light wave hitting a detector is an event.

Events can be observed and measured by experimenters in different reference frames. An exploding firecracker is as clear to you as you drive by in your car as it is to me standing on the street corner. We can quantify where and when an event occurs with four numbers: the coordinates (x, y, z) and the instant of time t . These four numbers, illustrated in **FIGURE 37.12**, are called the **spacetime coordinates** of the event.

The spatial coordinates of an event measured in reference frames S and S' may differ. It now appears that the instant of time recorded in S and S' may also differ. Thus the spacetime coordinates of an event measured by experimenters in frame S are (x, y, z, t) and the spacetime coordinates of the *same event* measured by experimenters in frame S' are (x', y', z', t') .

The motion of a particle can be described as a sequence of two or more events. We introduced this idea in the preceding section when we agreed to measure the velocity of a bicycle and then of a light wave by comparing the object passing the tree (first event) to the object passing the lamppost (second event).

Measurements

Events are what “really happen,” but how do we learn about an event? That is, how do the experimenters in a reference frame determine the spacetime coordinates of an event? This is a problem of *measurement*.

We defined a reference frame to be a coordinate system in which experimenters can make position and time measurements. That’s a good start, but now we need to be more precise as to *how* the measurements are made. Imagine that a reference frame is filled with a cubic lattice of meter sticks, as shown in **FIGURE 37.13**. At every intersection is a clock, and all the clocks in a reference frame are *synchronized*. We’ll return in a moment to consider how to synchronize the clocks, but assume for the moment it can be done.

Now, with our meter sticks and clocks in place, we can use a two-part measurement scheme:

- The (x, y, z) coordinates of an event are determined by the intersection of the meter sticks closest to the event.
- The event’s time t is the time displayed on the clock nearest the event.

You can imagine, if you wish, that each event is accompanied by a flash of light to illuminate the face of the nearest clock and make its reading known.

Several important issues need to be noted:

1. The clocks and meter sticks in each reference frame are imaginary, so they have no difficulty passing through each other.
2. Measurements of position and time made in one reference frame must use only the clocks and meter sticks in that reference frame.
3. There’s nothing special about the sticks being 1 m long and the clocks 1 m apart. The lattice spacing can be altered to achieve whatever level of measurement accuracy is desired.

FIGURE 37.12 The location and time of an event are described by its spacetime coordinates.

An event has spacetime coordinates (x, y, z, t) in frame S and different spacetime coordinates (x', y', z', t') in frame S' .

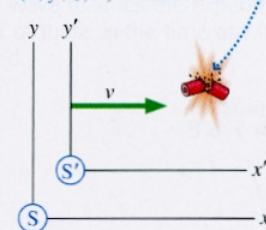
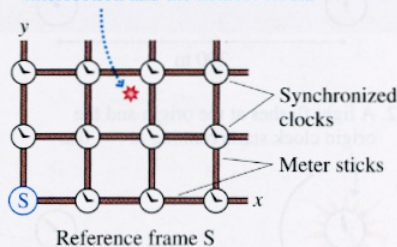
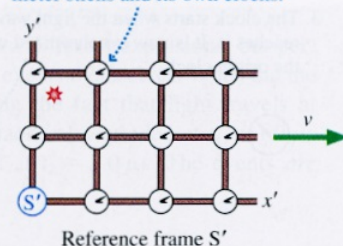


FIGURE 37.13 The spacetime coordinates of an event are measured by a lattice of meter sticks and clocks.

The spacetime coordinates of this event are measured by the nearest meter stick intersection and the nearest clock.



Reference frame S' has its own meter sticks and its own clocks.



4. We'll assume that the experimenters in each reference frame have assistants sitting beside every clock to record the position and time of nearby events.
5. Perhaps most important, t is the time at which the event *actually happens*, not the time at which an experimenter sees the event or at which information about the event reaches an experimenter.
6. All experimenters in one reference frame agree on the spacetime coordinates of an event. In other words, **an event has a unique set of spacetime coordinates in each reference frame.**

STOP TO THINK 37.3

A carpenter is working on a house two blocks away. You notice a slight delay between seeing the carpenter's hammer hit the nail and hearing the blow. At what time does the event "hammer hits nail" occur?

- a. At the instant you hear the blow
- b. At the instant you see the hammer hit
- c. Very slightly before you see the hammer hit
- d. Very slightly after you see the hammer hit

Clock Synchronization

It's important that all the clocks in a reference frame be **synchronized**, meaning that all clocks in the reference frame have the same reading at any one instant of time. Thus we need a method of synchronization. One idea that comes to mind is to designate the clock at the origin as the *master clock*. We could then carry this clock around to every clock in the lattice, adjust that clock to match the master clock, and finally return the master clock to the origin.

This would be a perfectly good method of clock synchronization in Newtonian mechanics, where time flows along smoothly, the same for everyone. But we've been driven to reexamine the nature of time by the possibility that time is different in reference frames moving relative to each other. Because the master clock would *move*, we cannot assume that the moving master clock would keep time in the same way as the stationary clocks.

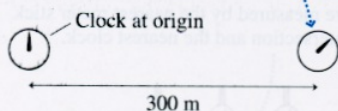
We need a synchronization method that does not require moving the clocks. Fortunately, such a method is easy to devise. Each clock is resting at the intersection of meter sticks, so by looking at the meter sticks, the assistant knows, or can calculate, exactly how far each clock is from the origin. Once the distance is known, the assistant can calculate exactly how long a light wave will take to travel from the origin to each clock. For example, light will take $1.00\ \mu\text{s}$ to travel to a clock $300\ \text{m}$ from the origin.

NOTE ▶ It's handy for many relativity problems to know that the speed of light is $c = 300\ \text{m}/\mu\text{s}$. ◀

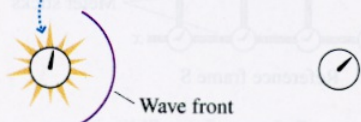
To synchronize the clocks, the assistants begin by setting each clock to display the light travel time from the origin, but they don't start the clocks. Next, as **FIGURE 37.14** shows, a light flashes at the origin and, simultaneously, the clock at the origin starts running from $t = 0\ \text{s}$. The light wave spreads out in all directions at speed c . A photodetector on each clock recognizes the arrival of the light wave and, without delay, starts the clock. The clock had been preset with the light travel time, so each clock as it starts reads exactly the same as the clock at the origin. Thus all the clocks will be synchronized after the light wave has passed by.

FIGURE 37.14 Synchronizing clocks.

1. This clock is preset to $1.00\ \mu\text{s}$, the time it takes light to travel $300\ \text{m}$.



2. A light flashes at the origin and the origin clock starts running at $t = 0\ \text{s}$.



3. The clock starts when the light wave reaches it. It is now synchronized with the origin clock.



Events and Observations

We noted above that t is the time the event *actually happens*. This is an important point, one that bears further discussion. Light waves take time to travel. Messages, whether they're transmitted by light pulses, telephone, or courier on horseback, take time to be delivered. An experimenter *observes* an event, such as an exploding firecracker, only at a *later time* when light waves reach his or her eyes. But our interest is in the event itself, not the experimenter's observation of the event. The time at which the experimenter sees the event or receives information about the event is not when the event actually occurred.

Suppose at $t = 0$ s a firecracker explodes at $x = 300$ m. The flash of light from the firecracker will reach an experimenter at the origin at $t_1 = 1.0 \mu\text{s}$. The sound of the explosion will reach a sightless experimenter at $t_2 = 0.88$ s. Neither of these is the time t_{event} of the explosion, although the experimenter can work backward from these times, using known wave speeds, to determine t_{event} . In this example, the spacetime coordinates of the event—the explosion—are (300 m, 0 m, 0 m, 0 s).

EXAMPLE 37.3 Finding the time of an event

Experimenter A in reference frame S stands at the origin looking in the positive x -direction. Experimenter B stands at $x = 900$ m looking in the negative x -direction. A firecracker explodes somewhere between them. Experimenter B sees the light flash at $t = 3.0 \mu\text{s}$. Experimenter A sees the light flash at $t = 4.0 \mu\text{s}$. What are the spacetime coordinates of the explosion?

MODEL Experimenters A and B are in the same reference frame and have synchronized clocks.

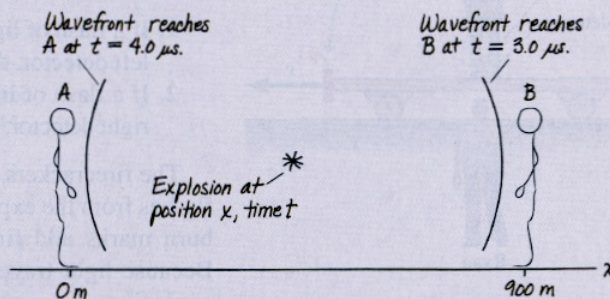
VISUALIZE FIGURE 37.15 shows the two experimenters and the explosion at unknown position x .

SOLVE The two experimenters observe light flashes at two different instants, but there's only one event. Light travels $300 \text{ m}/\mu\text{s}$, so the additional $1.0 \mu\text{s}$ needed for the light to reach experimenter A implies that distance $(x - 0 \text{ m})$ is 300 m longer than distance $(900 \text{ m} - x)$. That is,

$$(x - 0 \text{ m}) = (900 \text{ m} - x) + 300 \text{ m}$$

This is easily solved to give $x = 600 \text{ m}$ as the position coordinate of the explosion. The light takes $1.0 \mu\text{s}$ to travel 300 m to experi-

FIGURE 37.15 The light wave reaches the experimenters at different times. Neither of these is the time at which the event actually happened.



menter B, $2.0 \mu\text{s}$ to travel 600 m to experimenter A. The light is received at $3.0 \mu\text{s}$ and $4.0 \mu\text{s}$, respectively; hence it was emitted by the explosion at $t = 2.0 \mu\text{s}$. The spacetime coordinates of the explosion are $(600 \text{ m}, 0 \text{ m}, 0 \text{ m}, 2.0 \mu\text{s})$.

ASSESS Although the experimenters *see* the explosion at different times, they agree that the explosion *actually happened* at $t = 2.0 \mu\text{s}$.

Simultaneity

Two events 1 and 2 that take place at different positions x_1 and x_2 but at the *same time* $t_1 = t_2$, as measured in some reference frame, are said to be **simultaneous** in that reference frame. Simultaneity is determined by when the events *actually* happen, not when they are seen or observed. In general, simultaneous events are *not* seen at the same time because of the difference in light travel times from the events to an experimenter.

EXAMPLE 37.4 Are the explosions simultaneous?

An experimenter in reference frame S stands at the origin looking in the positive x -direction. At $t = 3.0 \mu\text{s}$ she sees firecracker 1 explode at $x = 600 \text{ m}$. A short time later, at $t = 5.0 \mu\text{s}$, she sees firecracker 2 explode at $x = 1200 \text{ m}$. Are the two explosions simultaneous? If not, which firecracker exploded first?

MODEL Light from both explosions travels toward the experimenter at $300 \text{ m}/\mu\text{s}$.

SOLVE The experimenter *sees* two different explosions, but perceptions of the events are not the events themselves. When did the explosions *actually* occur? Using the fact that light travels at $300 \text{ m}/\mu\text{s}$, we can see that firecracker 1 exploded at $t_1 = 1.0 \mu\text{s}$ and firecracker 2 also exploded at $t_2 = 1.0 \mu\text{s}$. The events *are* simultaneous.